An Introduction to Surface Wave Analysis

ADVANCED STUDIES COURSE IN JOINT INVERSION OF RECEIVER FUNCTIONS AND SURFACE WAVE DISPERSION Kuwait City, KUWAIT 21 January, 2013



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Outline

Section 1 – Background on Surface Waves - Basics, Applications, Results, Mideast

Section 2 – Surface Wave Analysis – Measurement, Tomography, Inversion

Section 3 – Model and Access Tool

Section 4 - Using Surface Waves to Build a Global Model - LITHO1.0 model The Basics of Surface Waves

- Surface Waves vs. Body Waves
- Group Velocity vs. Phase Velocity
- Rayleigh Waves vs. Love Waves
- Higher Mode Surface Waves
- Historical Background
- Variations due to crustal thickness, sediments thickness, etc.
- Sensitivity Kernels

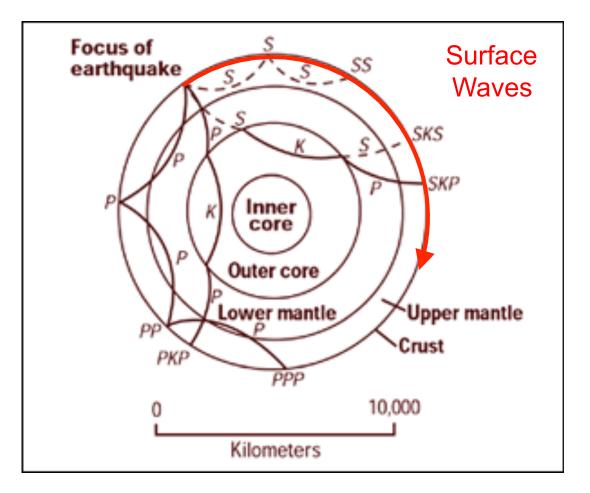
Applications

Results

Examples from the Middle East

Surface Waves

Surface waves refer to seismic waves that travel along the earth's surface, as opposed to *body waves*, which travel through the earth's interior



Surface waves are *dispersive*, which gives the characteristic that their velocities are a function of frequency (period) and that their group velocities are generally not equal to their phase velocities.

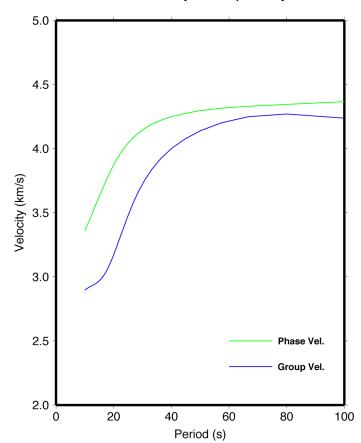
Group Velocity (U = $d\omega/dk$) is the velocity in which the wave energy moves

Phase Velocity (C = ω/k) is the velocity that a peak or trough moves

red dot = phase velocity green dot = group velocity

Group Velocity vs. Phase Velocity

- For typical earth profiles, the phase velocity is generally faster than the group velocity.
- While the group velocity can increase or decrease with increasing period, phase velocities are monotonically increasing.
- If the group velocities are constant over a wide period range, then they can produce a high-amplitude body-wave looking pulse that is called an *Airy phase*



Phase Velocity vs. Group Velocity

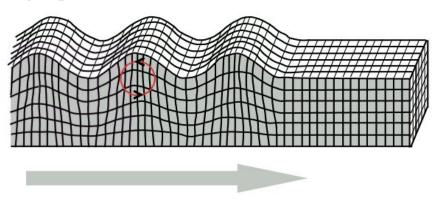
Rayleigh Waves

Designated LR = "Long-period Rayleigh" Also referred to as "ground roll" in refraction/reflection surveys and other applications

They are sensitive to P-SV

Produce retrograde motion on Z (vertical) and R (radial) components

Rayleigh Wave

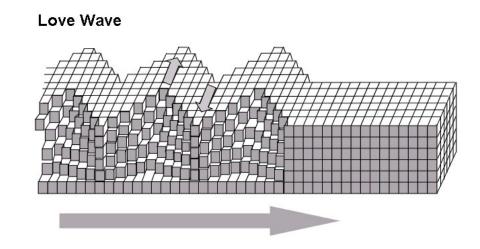


Love Waves

Designated LQ = "Long-period Querwellen" (German for transverse waves)

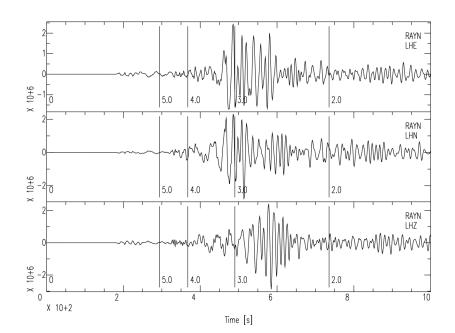
Sensitive to SH (horizontally propagating shear waves)

Shear motion on T (transverse) component



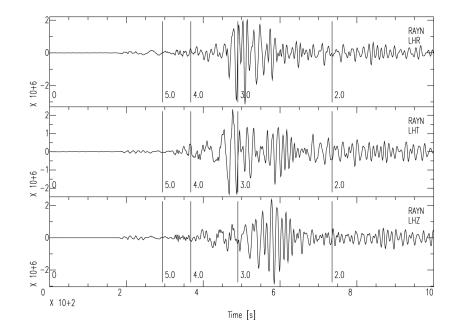
Rayleigh Waves vs. Love Waves

Unrotated – LHE,LHN,LHZ



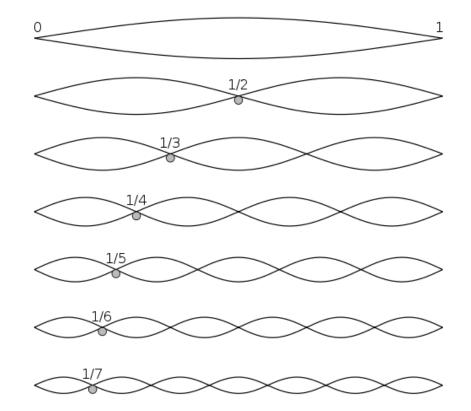
Rotated – LHR,LHT,LHZ





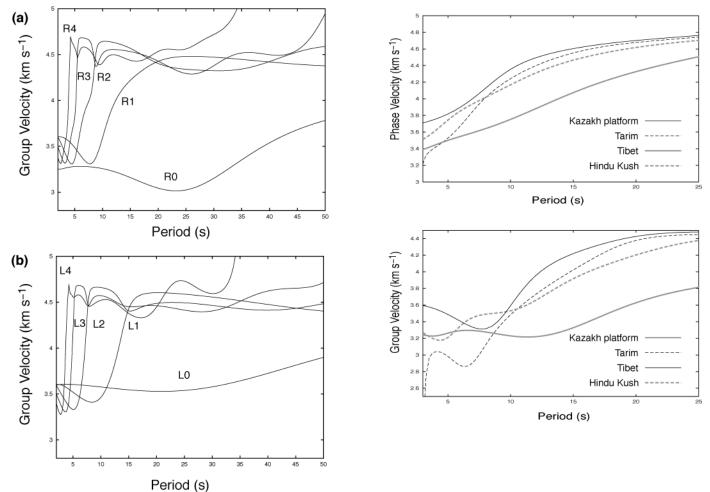
So far, we have been talking about *fundamental mode* surface waves.

There exists an infinite number of higher mode solutions



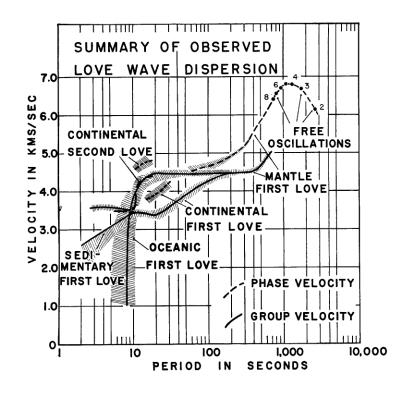
Higher Modes

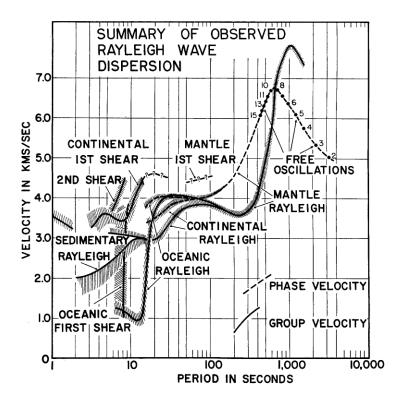
Higher mode exist, but are more difficult to measure



From Levshin et al. (2005)

Surface wave studies have been around for a long time in seismology



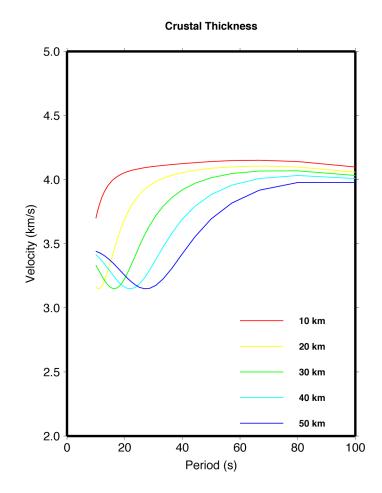


From Oliver (1962)

Dispersion Variations: Crustal Thickness

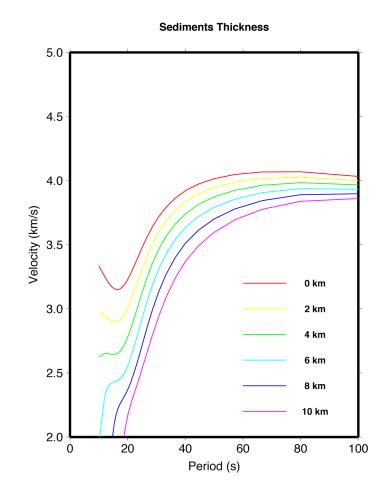
Oceanic crust typically has crustal thicknesses of < 10 km

Continental crust has crustal thickness that ranges between 15 and 80 km, but typically around 35 km



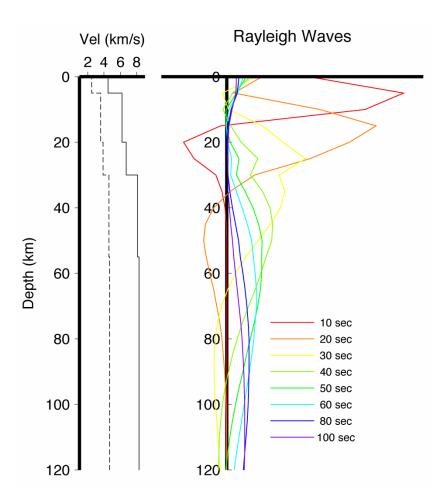
Dispersion Variations: Sediments Thickness

Sediment thicknesses range from 0 km to over 20 km.



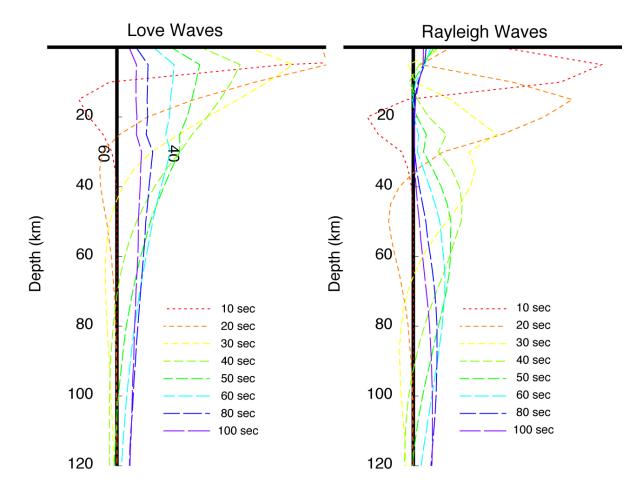
Sensitivity kernels show the relationship between dispersion velocities and earth structure.

- They are calculated by taking the partial derivative of the dispersion velocities with respect to other parameters, such as shear-wave velocity. For instance, $\delta U/\delta \beta$ is often shown here.
- With increasing period, surface waves become sensitive to deeper velocity structures.
- They are themselves dependent on the velocity structure, so it is non-linear.
- Sensitivity kernels are used to invert the dispersion curves for layered velocity structure.



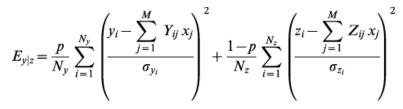
Rayleigh waves and Love waves have different sensitivity kernels

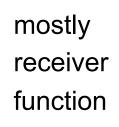
The longer the period, the deeper the structure sampled

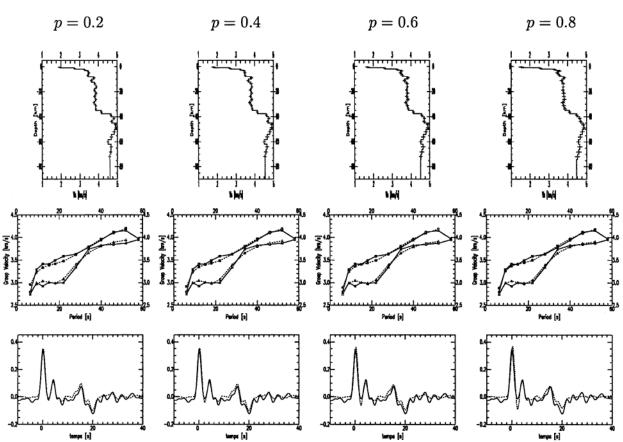


Applications

Joint inversion of surface waves and receiver functions (Julia et al., 2000)





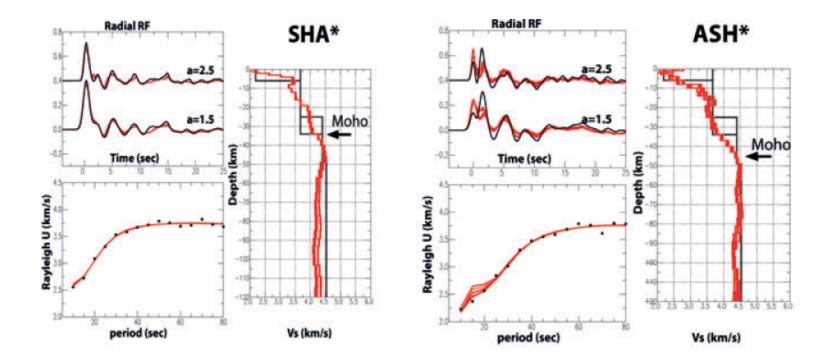


mostly surface waves

Applications

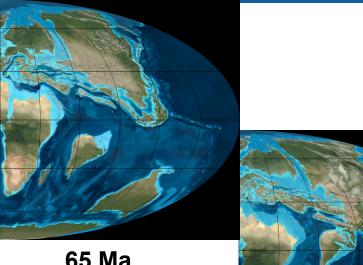
Example of joint inversion results in Oman (Al-Hashmi et al., 2011)

- red lines show results from different starting models, influence parameters (p=0.3,0.5,0.7), and layering smoothness to show range



Al-Hashmi, S., R. Gök, K. Al-Toubi, Y. Al-Shijbi, I. El-Hussain, and A.J. Rodgers (2011). Seismic velocity structure at the southeastern margin of the Arabian Peninsula, Geophys. J. Int., DOI: 10.1111/j.1365-246X.2011.05067.x

Recent Tectonic History of Middle East Region



65 Ma

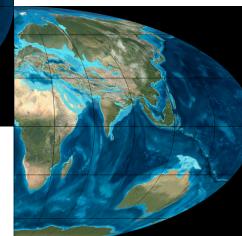


- assembly of terranes in southern Eurasia prior to the collision of Africa and India

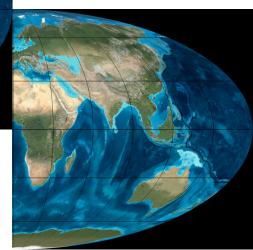
- orogeny and plateau building along Tethys collision zone

- continuing subduction in the Mediterranean and in the Makran

35 Ma







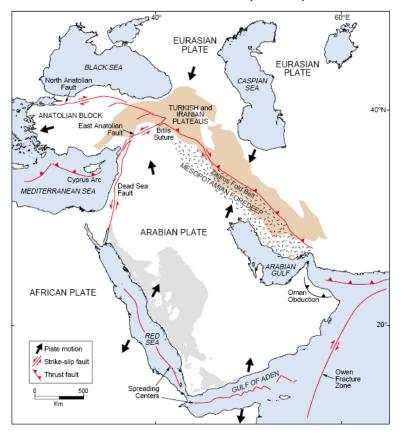
Figures courtesy of Ron Blakey, Northern Arizona University

50 Ma

Quick tectonic overview

- Convergence between Arabian and Eurasia Plates producing the continued uplift of the Zagros Mts. and Turkish and Iranian Plateaus
- Rifting along axes of Afar Triple Junction creating Red Sea, Gulf of Aden, and East African Rift Zone
- Remnant oceanic crust in Black and Caspian Seas
- Subduction of oceanic crust in the Mediterranean Sea and along the Makran
- Large strike-slip faults along the Dead Sea Fault, East Anatolian Fault, and North Anatolian Fault
- Very deep sedimentary basins along the Persian/Arabian Gulf, Mesopotamian Foredeep, eastern Mediterranean and Caspian Sea.
- Precambrian Arabian-Nubian Shield

Simplified tectonic map from Seber et al. (2000)



Surface Wave Results for the Middle East

Model from Ma et al. (2012) in preparation

Data Rayleigh Wave Group Velocity 10.00 mHz

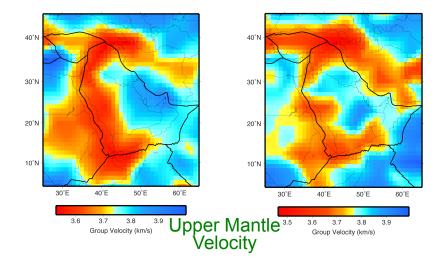
40°N

30°N

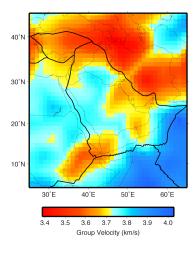
20°N

10°N

Data Rayleigh Wave Group Velocity 15.00 mHz

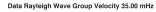


Data Rayleigh Wave Group Velocity 20.00 mHz

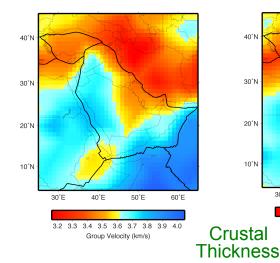


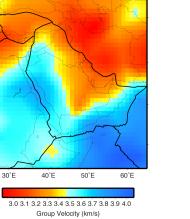
Data Rayleigh Wave Group Velocity 25.00 mHz

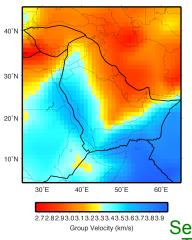
Data Rayleigh Wave Group Velocity 30.00 mHz

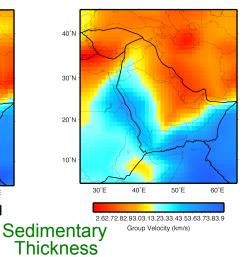


Data Rayleigh Wave Group Velocity 40.00 mHz









Summary

- Surface waves are a well-known and well-understood way of studying the earth
- Surface waves are effective at sampling aseismic regions.
- The differing sensitivities of Love and Rayleigh wave phases, and for different periods, allows us to sample the depth profile of the earth.
- We will use them in conjunction with receiver functions to develop more robust models.

Section 2 - Surface Wave Analysis – Measurement and Tomography

Surface Wave Measurements (Dispersion along Paths)

• Event Based, Ambient Noise, Clustering

Seismic Tomography (Lateral Dispersion)

• System of equations, Damping

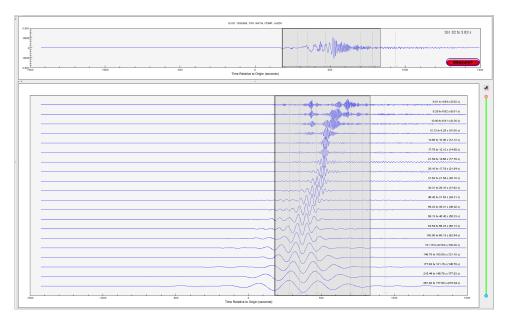
Layered Velocity Inversion (Dispersion to Velocity)

Inversion vs. Grid Search

Multiple-Filter Analysis

Multiple-Filter Analysis – a narrow-band Gaussian filter is applied over many different periods (e.g. Dziewonski, et al., 1969; Herrmann, 1973)

FTAN (Frequency-Time Analysis) – Levshin et al., 1972

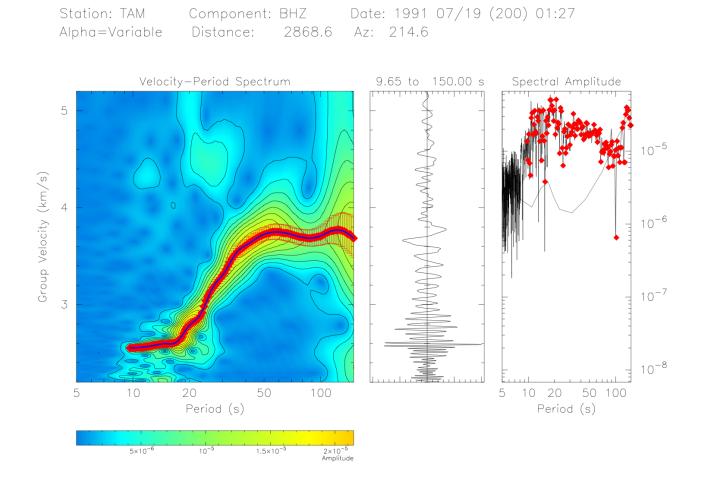


Dziewonski, A., S. Bloch, and M. Landisman (1969). A technique for the analysis of transient seismic signals, Bull. Seism. Soc. Amer., 59, 427-444.

Herrmann, R.B. (1973). Some aspects of bandpass filtering of surface waves, Bull. Seism. Soc. Amer., 63, 663-671.

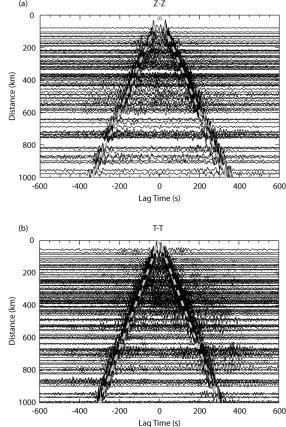
Levshin, A.L., V.F. Pisarenk, G.A. Pogrebin (1972), Frequency-time analysis of oscillations, Annales de Geophysique, 28, 211

PGSWMFA = PGplot Surface Wave Multiple Filter Analysis uses PGPLOT plotting package



Ambient Noise Analysis

Algorithms developed by the University of Colorado group (and detailed in Bensen et al., 2007)

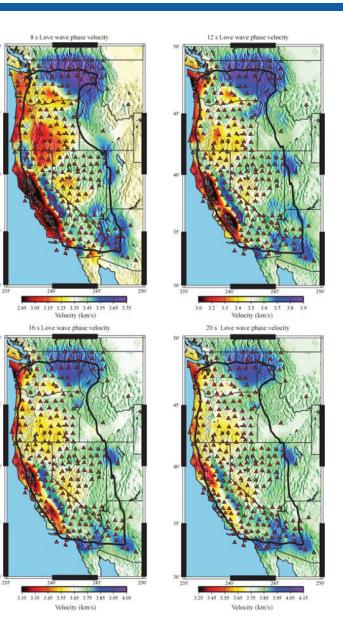


Bensen, G.D., M.H. Ritzwoller, M.P. Barmin, A.L. Levshin, F. Lin, M.P. Moschetti, N.M. Shapiro, and Y. Yang (2007), Processing seismic ambient noise data to obtain reliable broad-band surface wave dispersion measurements, Geophys. J. Int., (169), 1239–1260.

Ambient Noise Analysis

Has the potential to produce very high resolution models where station coverage is dense

Lin, F.-C., M.P. Moschetti, and M.H. Ritzwoller (2008). Surface wave tomography of the western United States from ambient seismic noise: Rayleigh and Love wave phase velocity maps, Geophys. J. Int. doi: 10.1111/j.1365-246X.2008.03720.x



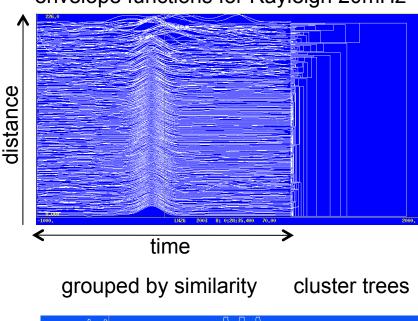
Surface wave model developed using cluster analysis

We have developed a new, very large global surface wave dataset build using a new, efficient measurement technique that employs cluster analysis

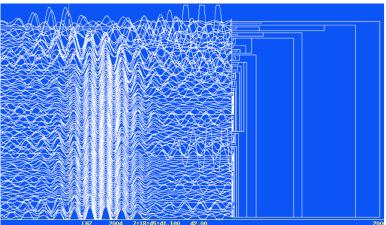


Measurement algorithm

- "Undisperse" using a 1-D phase velocity curve
- Correct for source phase and amplitude
- Correct for predicted phase shift from 3D structure using a nominal phase velocity map



envelope functions for Rayleigh 20mHz



Ma et al. (2012) in preparation

Rayleigh 10mHz

Consistency between group and phase velocity maps is achieved through cubic splines

Relationship between group velocity U and phase velocity c

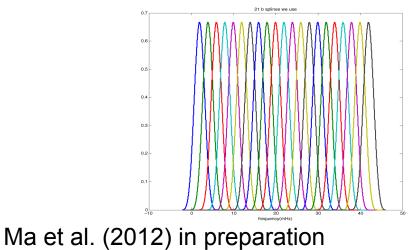
$$\frac{1}{U} = \frac{1}{c} + \omega \frac{d}{d\omega} \frac{1}{d\omega}$$

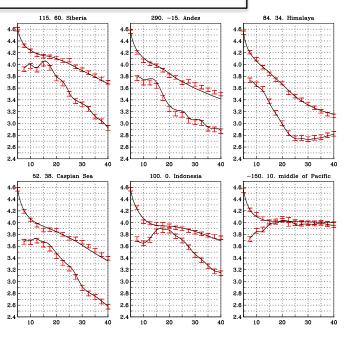
or

$$\frac{1}{U} = \sum B_i(\omega)a_i$$
$$\frac{\omega}{c} = \sum \left[\int_{\omega_0}^{\omega} B(\omega')d\omega'\right]a_i + a_0$$

$$\frac{1}{c} = \sum B_i(\omega)a_i$$
$$\frac{1}{U} = \sum [\omega B'_i(\omega) + B_i(\omega)]a_i$$

where $B_i(\omega)$ are b spline functions, $B_i'(\omega)$ are the derivatives and a_i are the coefficients



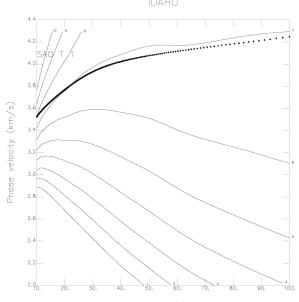


Single station (event-station)

- Need to know the source phase ϕ_0 , 2π indeterminacy

Mechanism and depth needed to determine source phase ϕ_0 .

Unwrapping the phase by working from long periods.



AMPLITUDE PHASE
$$u(x,t) = \frac{1}{\pi} \int_{0}^{\infty} \hat{u}(\omega, x) \times \cos\left(\omega t - \frac{\omega}{c(\omega)} x + \phi_{0}(\omega)\right) d\omega$$

 $\phi(\omega) = \phi_0(\omega) - \frac{\omega x_1}{c(\omega)} + 2\pi N + \omega t$

$$\psi_1(\omega) = \omega t_1 + \phi_o(\omega) - \frac{\omega x_1}{c(\omega)} + 2\pi N$$



Two-station method (event-event)

- Source phase ϕ_0 cancels
- Still have 2π indeterminacy

Two-plane method (Forsyth, 1998)

 uses the sum of two plane waves, each with initially unknown amplitude, initial phase, and propagation direction to represent the nonplanar incoming wavefield, i.e., a total of six parameters to describe the incoming wavefield.

$$\psi_1(\omega) - \psi_2(\omega) = \omega(t_1 - t_2) - \frac{\omega}{c(\omega)}(x_1 - x_2) + 2\pi M$$

$$c(\omega) = \frac{x_1 - x_2}{(t_1 - t_2) + T[M - (1/2\pi)(\psi_1(\omega) - \psi_2(\omega))]}$$

$$U(\omega) = A_1(\omega) \exp(-i\phi_1) + A_2(\omega) \exp(-i\phi_2)$$

$$\phi_1 = \phi_1^0 + \omega[r\cos(\psi - \theta_1) - x]/c(\omega) + \omega(\tau - \tau_0)$$

$$\phi_2 = \phi_2^0 + \omega[r\cos(\psi - \theta_2) - x]/c(\omega) + \omega(\tau - \tau_0)$$

Tomography is an imaging method widely used in seismology for the derivation of bulk earth properties such as velocity and attenuation from measured properties along paths.

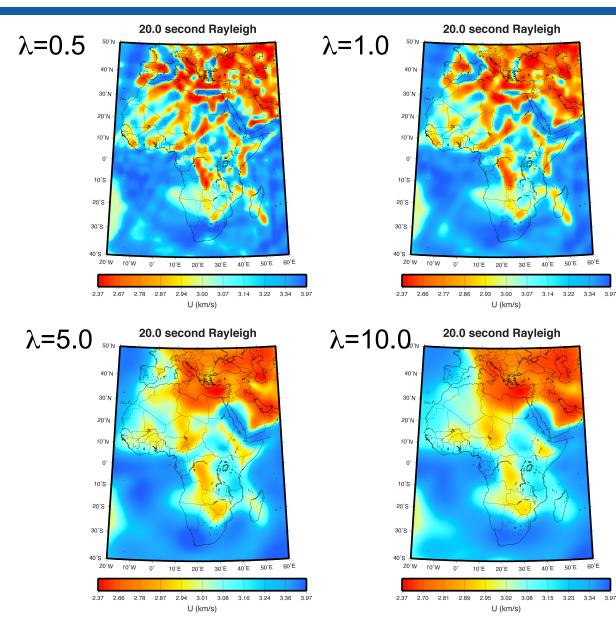
It is analogous to tomography methods used in medical imaging, such as CT scans, but generally with much poorer coverage of the study region.

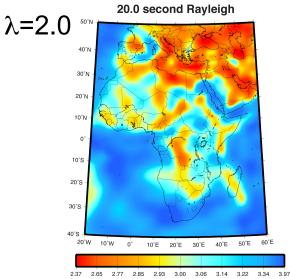
t = A x

We use it to invert dispersion measurements for spatially varying dispersion values.

There are many methods to solve the inversion. We will be using a program that uses the conjugate gradient method.

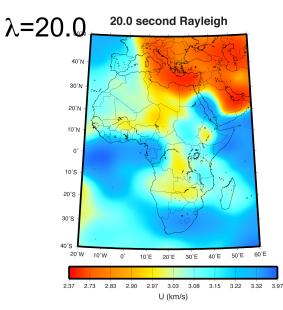
Seismic Tomography – Importance of Damping





U (km/s)

3.97



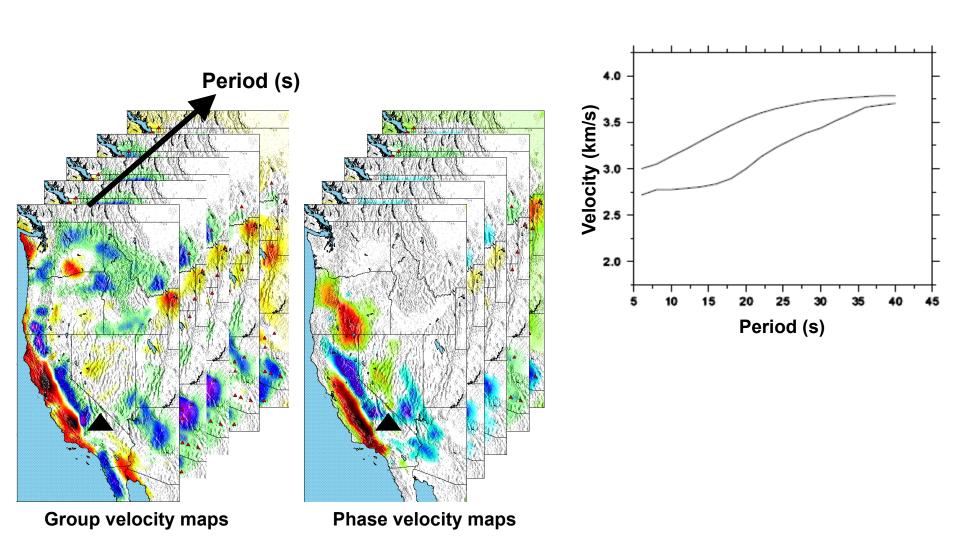
Inverting group velocities and phase velocities for layered earth structure is both:

- Non-unique (many possible models can fit the same dispersion data)
- Non-linear (the sensitivity kernels used to invert for the earth structure itself depend on the earth structure)

I usually employ a grid-search method to estimate the layered velocity structure.

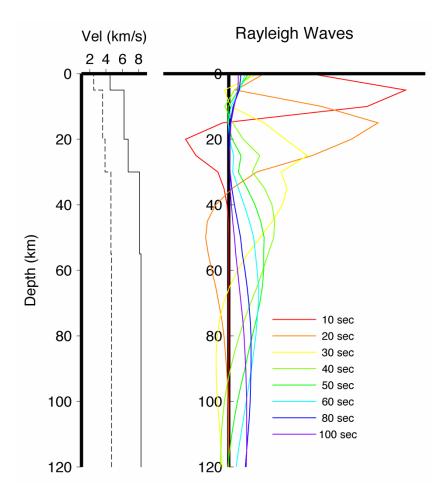
One advantage is that one doesn't actually perform an inversion, just a series of forward calculations. This allows us to explore the model space.

Getting from dispersion maps and dispersion curves to layered velocity structure



Sensitivity kernels show the relationship between dispersion velocities and earth structure.

- They are calculated by taking the partial derivative of the dispersion velocities with respect to other parameters, such as shear-wave velocity. For instance, $\delta U/\delta \beta$ is often shown here.
- With increasing period, surface waves become sensitive to deeper velocity structures.
- They are themselves dependent on the velocity structure, so it is non-linear.
- We will be using them later when we invert our dispersion curves for layered velocity structure.



The inverse problem can be conceptually formulated as follows:

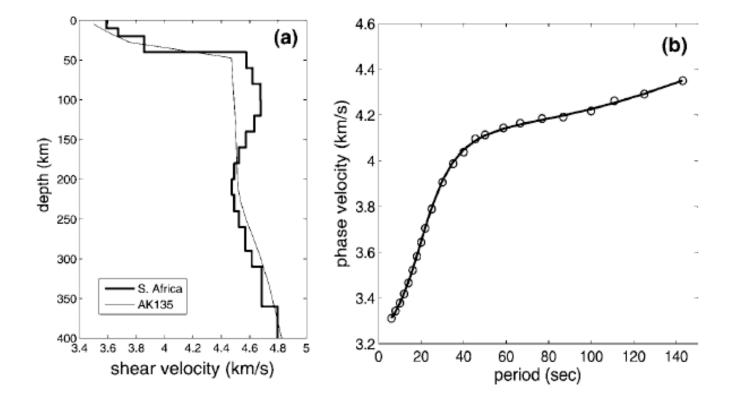
Data \rightarrow **Model** parameters

The inverse problem is considered the "inverse" to the forward problem which relates the model parameters to the observed data:

Model parameters → Data

A grid search performs an inversion by calculating the forward problem multiple times and comparing the predicted data to the observed data.

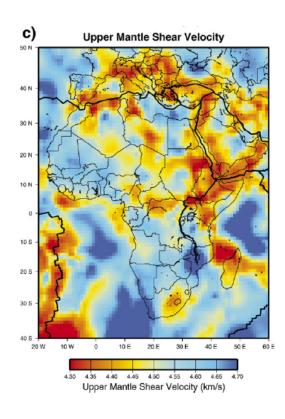
Example of inverting surface wave dispersion maps for 1-D shear-velocity profile



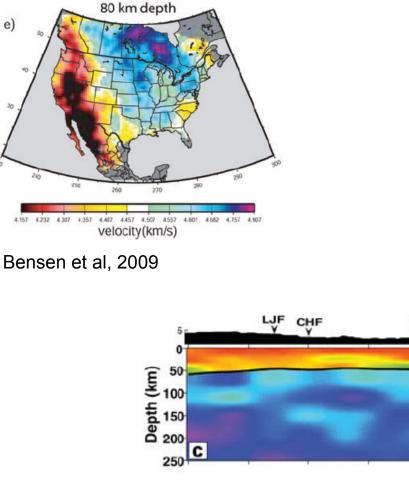
Results can be combined to produce 3-D structural models

e)

Examples of structure derived from surface waves



Pasyanos and Nyblade, 2007



Yao et al, 2008

PDHF

LZJ

Inversion vs. Grid-Search

Inversion

Issues:

- Invert for layer thickness
- Invert for layer velocity
- Starting model
- Non-linear inversion
- Layer smoothing

Grid Search

Issues:

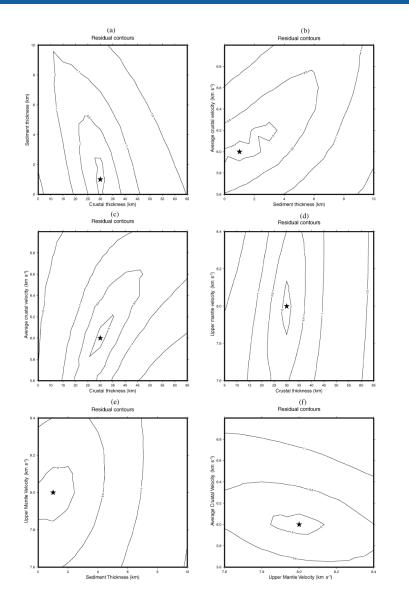
- Fewer inversion parameters
- Usually slower
- Unique dispersion curve for each trial model
- Can be used to map out model space

Programs

Computer Programs in Seismology (Bob Herrmann, SLU) is a nice package of software to enable inversion of dispersion measurements (and a lot more)

MINOS (Woodhouse, Masters) is a comprehensive package that computes surface wave dispersion by summing normal modes

Tradeoffs among various parameters



Pasyanos, M.E. and W.R. Walter (2002). Crust and upper-mantle structure of North Africa, Europe and the Middle East from inversion of surface waves, Geophys. J. Int. 149, 463-481.

Adding other types of data (for example, receiver functions) can help reduce these tradeoffs

Inversion Parameters

Possible inversion parameters

- Sediment thickness
- Crustal thickness
- Crustal velocity
- Crustal Vp/Vs
- Upper mantle velocity
- Upper mantle Vp/Vs
- Upper mantle anisotropy
- Lithospheric thickness

One approach employed in Pasyanos and Nyblade (2007) and Pasyanos (2010)

Grid search

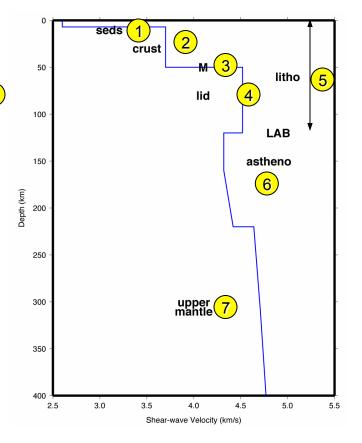
- fix sediments from Laske sediment profile ¹
- solve for v_p/v_s/2 crustal thickness, 3p_n/s_n, 4 lithospheric thickness 5
- asthenosphere⁶ has lower Vp and higher Poisson's ratio (σ =0.29)
- upper mantle⁷ is transitioned into ak135 model (Kennett et al., 1995)

NOT solving for

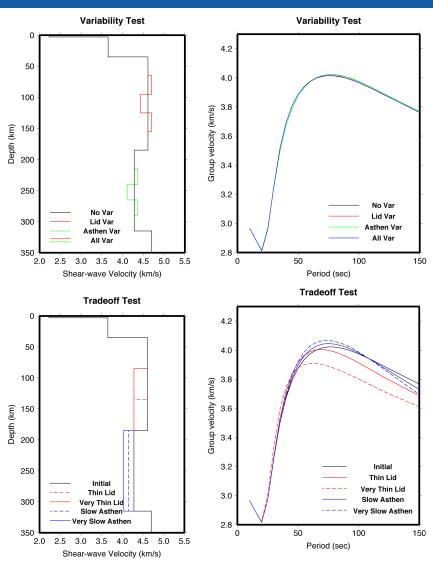
detailed variations in velocity or Poisson's ratio in crust or lid

Pasyanos, M.E. and A.A. Nyblade (2007), A top to bottom lithospheric study of Africa and Arabia, Tectonophysics, 444, 27-44, doi:10.1016/j.tecto. 2007.07.008.

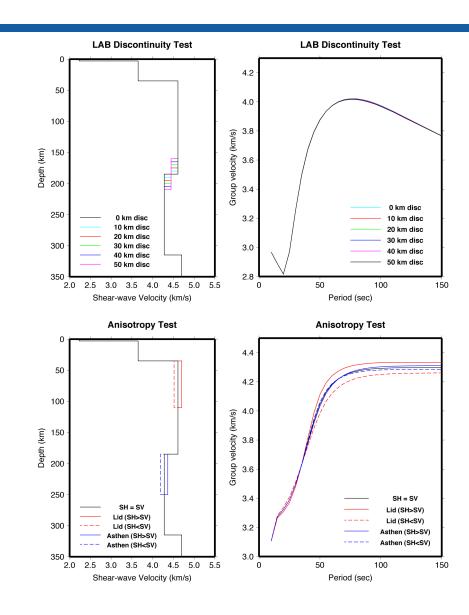
Pasyanos, M.E. (2010). Lithospheric thickness modeled from long-period surface wave dispersion, Tectonophys., 481, 38-50.



Inversion Tests







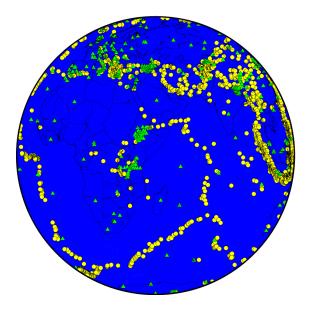
Summary

- The analysis of seismic surface waves are a well-established method of estimating earth structure.
- There are several ways of measuring seismic dispersion
- Seismic tomography can be used to invert those values for lateral variations in dispersion
- Another inversion must be used to determine what velocity structure is consistent with the observed dispersion
- Profiles over broad regions can be combined to produce 3-D structural models.

We have provided a tool to retrieve dispersion information from the global surface wave model of Ma et al. (2012).

Surface Wave Model - Coverage

Surface Wave Paths



Paths at each frequency

Frequency	/ (mHz)	5	7.5	10	12.5	15	17.5	20	22.5	25	27.5	30	32.5	35	37.5	40
Rayleigh	Phase	305k	502k	582k	609k	603k	595k	400k	525k	414k	489k	320k	403k	282k	()	()
	Group		316k	353k	331k	299k	316k	334k	322k	199k	210k	203k	214k	109k	110k	113k
Love	Phase		140k	198k	246k	175k	222k	145k	219k	152k	194k	108k				
	Group			189k	182k	166k	180k	171k	176k	166k	77k	76k				

Surface Wave Model - Results

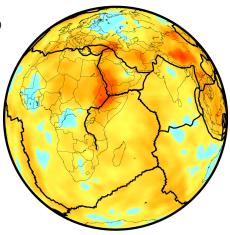
Rayleigh Wave Group Velocity 30.0 mHz

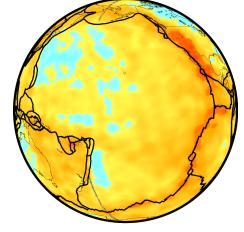
Rayleigh Wave Group Velocity 30.0 mHz

 3.0
 3.1
 3.2
 3.3
 3.4
 3.5
 3.6
 3.7
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 3.9
 4.03.0
 3.1
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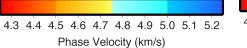
 Group Velocity (km/s)

Love Wave Phase Velocity 10.0 mHz





Love Wave Phase Velocity 10.0 mHz



4.3 4.4 4.5 4.6 4.7 4.8 4.9 5.0 5.1 5.2 Phase Velocity (km/s)

Examples

get_dispersion 23.522499 45.503201

Getting dispersion values from surface wave model lat = 23.5224990000000 lon = 45.503201000000 avg nblk= 1 15 nfreg = 10 RAYLEIGH Freq (mHz) Period (s) Group Vel (km/s) Phase Vel (km/s) 5.0000 200.0000 -99.0000 0.0400 4.5560 0.0300 7.5000 133.3333 3.6787 0.0400 4.1909 0.0300 10.0000 100.0000 3.7663 0.0400 4.0654 0.0300 12,5000 80,0000 3,7896 4.0049 0.0300 0.0400 15.0000 66.6667 3.8010 0.0400 3.9687 0.0300 17.5000 57.1429 3.7885 0.0400 3.9434 0.0300 20.0000 50.0000 3.6876 0.0400 3.9173 0.0300 22.5000 44.4444 3.6096 0.0400 3.8846 0.0300 25.0000 40.0000 3.5748 0.0400 3.8535 0.0300 27.5000 36.3636 3.4627 3.8212 0.0300 0.0400 30.0000 33.3333 3.3749 0.0400 3.7834 0.0300 32,5000 30,7692 3,3130 0.0400 3.7456 0.0300 35.0000 28.5714 3.2361 3.7075 0.0300 0.0400 37.5000 26.6667 3.1533 0.0400 3.6684 0.0300 40.0000 25.0000 3.0424 0.0400 3.6269 0.0300 LOVE Freq (mHz) Period (s) Group Vel (km/s) Phase Vel (km/s) 7.5000 133.3333 -99.0000 0.0600 4.6965 0.0400 4.5931 10.0000 100.0000 4.2473 0.0600 0.0400 12.5000 80.0000 4.0981 0.0600 4.5036 0.0400 15.0000 66.6667 4.0472 0.0600 4.4231 0.0400 17.5000 57.1429 3.9942 0.0600 4.3620 0.0400 20.0000 50.0000 3.9145 0.0600 4.3057 0.0400 22,5000 44,4444 0.0600 4.2554 3.8609 0.0400 25.0000 40.0000 3.7287 0.0600 4.2044 0.0400 27.5000 36.3636 3.6923 0.0600 4.1535 0.0400 0.0600 30.0000 33.3333 3.7505 4.1119 0.0400

get_dispersion 29.17556 47.69333

Getting dispersion values from surface wave model lat = 29.175560000000 lon = 47.693330000000 avg nblk= 1 15 10 nfreg = RAYLEIGH Freq (mHz) Period (s) Group Vel (km/s) Phase Vel (km/s) 5.0000 200.0000 -99.0000 0.0400 4.5539 0.0300 7.5000 133.3333 3.8017 0.0400 4.2537 0.0300 4.1443 0.0300 10.0000 100.0000 3.9035 0.0400 12,5000 80,0000 3,9397 0.0400 4.1014 0.0300 0.0400 15.0000 66.6667 3.8425 4.0644 0.0300 17.5000 57.1429 3.7577 0.0400 4.0246 0.0300 20.0000 50.0000 3.7232 0.0400 3.9861 0.0300 22,5000 44,4444 3.6562 0.0400 3.9517 0.0300 25.0000 40.0000 3.5130 0.0400 3.9121 0.0300 27.5000 36.3636 3.3422 0.0400 3.8624 0.0300 30.0000 33.3333 3.1968 0.0400 3.8045 0.0300 32.5000 30.7692 3.0621 0.0400 3.7428 0.0300 35.0000 28.5714 2.8974 0.0400 3.6751 0.0300 37.5000 26.6667 2.7815 0.0400 3.6040 0.0300 40.0000 25.0000 2.6532 0.0400 3.5327 0.0300 LOVE Freq (mHz) Period (s) Group Vel (km/s) Phase Vel (km/s) 7.5000 133.3333 -99.0000 0.0600 4.7466 0.0400 10.0000 100.0000 4.4932 0.0600 4.6999 0.0400 12.5000 80.0000 4.2743 0.0600 4.6317 0.0400 15.0000 66.6667 4.0344 0.0600 4.5447 0.0400 17.5000 57.1429 3.8634 0.0600 4.4478 0.0400 20.0000 50.0000 3.7098 0.0600 4.3528 0.0400 22,5000 44,4444 3.5095 4.2558 0.0600 0.0400 25.0000 40.0000 0.0600 4.1586 3.4041 0.0400 27.5000 36.3636 3.2392 0.0600 4.0659 0.0400 3.9727 0.0400 30.0000 33.3333 3.1273 0.0600

get_dispersion is a tool which retrieves the surface wave dispersion at any given location

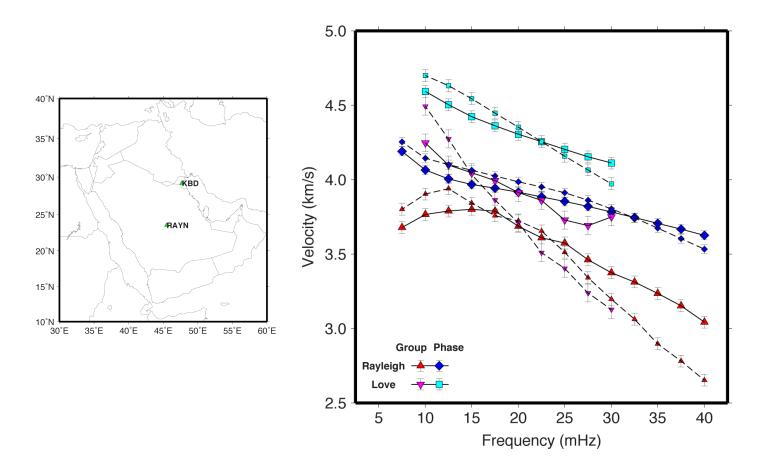
Surface wave data is from the model of Ma, Masters, Laske (UCSD Scripps) and Pasyanos (LLNL)

Rayleigh/Love group/phase 5 - 40mHz (200 – 25 sec period)

get_dispersion [lat] [lon] get_dispersion 23.522499 45.503201

Tools

Comparing between dispersion at KBD and RAYN.



Phase velocities somewhat similar, but group velocities, which sample shallower structure (especially at high frequencies) are very different.

We present an example from a current research project:

"LITHO1.0 – An updated crust and lithospheric model of the Earth developed using multiple data constraints" Michael E. Pasyanos (LLNL) Guy Masters, Gabi Laske and Zhitu Ma (IGPP, UC San Diego)

This section is a modified version of a talk that we gave at the Fall 2012 AGU Meeting in San Francisco

Introduction

Crustal models like CRUST5.1 (Mooney et al., 1998) and successor models like CRUST2.0 (Bassin et al., 2000) have been well-utilized in the seismological and geophysical communities (e.g. CRUST5.1 has 474 citations in Web of Science 11/2012)

Useful to varied sections of the communities

- global tomography in order to remove the crustal "noise" to see the mantle "signal"
- smaller scale crustal studies as starting model or comparison model

Downsides

- Poor predictor of travel times
- Resolution poor for many applications
- Model is of questionable quality in many poorly-covered regions

We are creating a 1 degree model of the crust and upper mantle that is consistent with

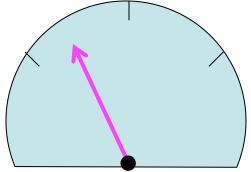
- Our extensive knowledge on sedimentary basins
- Improved information on crustal thickness
- Upper mantle velocities from regional travel times (e.g. Pn, Sn)
- High-resolution surface wave models across a broad frequency band

Our goal is to develop a higher-resolution model that extends deeper into the mantle to include the lithospheric lid. This is the LITHO1.0 model

Method

Our philosophy is to test a large series of models which are perturbations of a starting model which is consistent with other information about crust and upper mantle structure (e.g. tectonic regions, crustal thickness from receiver functions and other information, upper mantle velocities from travel time models, thermo-tectonic information, etc.)

There is a balance between honoring prior information and allowing enough variation to fit the surface wave data



Starting model

- Honors prior information
- Model may be poor
- Doesn't allow sufficient variations
- Often doesn't fit surface wave data

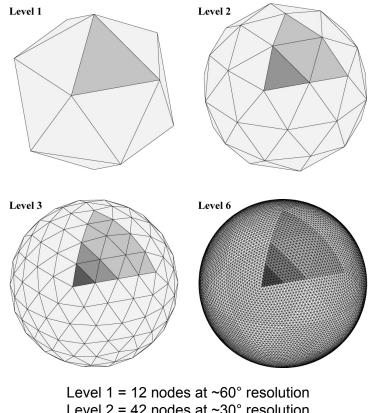
Surface Waves

- Provides varied models
- More sensitive to bulk properties, rather than discontinuities
- Doesn't honor prior information
- Important for global tomography models

Parameterization

Lateral Parameterization is achieved through tessellated nodes

Depth parameterization through the thickness and associated parameters of layers



Level 2 = 42 nodes at ~30° resolution Level 3 = 162 nodes at ~15° resolution Level 4 = 642 nodes at ~8° resolution Level 5 = 2562 nodes at ~4° resolution Level 6 = 10,242 nodes at ~2° resolution Level 7 = 40,962 nodes at ~1° resolution

Layer	Layer Name	Associated Parameters
W	Water/Ice	thick, vp, vs, density, Qp, Qs
S1	Sediment Layer 1	thick, vp, vs, density, Qp, Qs
S2	Sediment Layer 2	thick, vp, vs, density, Qp, Qs
S3	Sediment Layer 3	thick, vp, vs, density, Qp, Qs
C1	Upper Crust	thick, vp, vs, density, Qp, Qs
C2	Middle Crust	thick, vp, vs, density, Qp, Qs
C3	Lower Crust	thick, vp, vs, density, Qp, Qs
M1	Lithospheric Lid	thick, vp, vs, density, Qp, Qs
M2	Asthenosphere	thick, vp, vs, density, Qp, Qs
M3	Upper Mantle	thick, vp, vs, density, Qp, Qs

Building the starting model

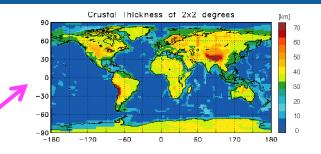
CRUST1.0 prototype

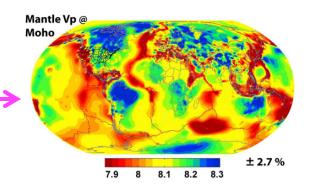
- Modified CRUST2.0 crustal model at higher resolution
- Full three-layer sediment model (Laske and Masters, 1997)
- Updated crustal thickness map

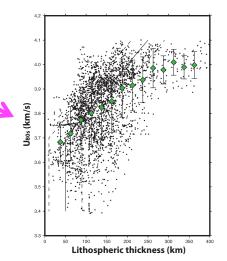
Upper mantle velocities from LLNL-G3Dv3 derived from regional and teleseismic travel times (Simmons et al., 2012)

Lithospheric thickness from the regression of long period dispersion and lithospheric thickness estimates from heat flow (continents) and lithospheric cooling (oceans) (Pasyanos, 2005)

Transitioned into ak135 model (Kennett et al., 1995) at depth



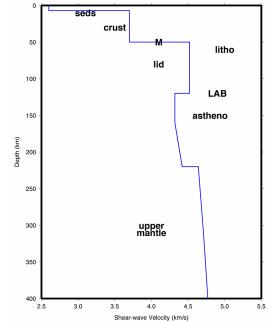




Varying the starting model

After creating an initial model, we perturb a number of model parameters (crustal velocities, mantle velocities, crustal thickness, lid thickness) to create a suite of about 10,000 models.

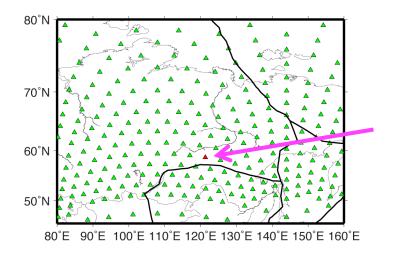
Water/ice/sediment structure – fixed Crustal velocity stack – perturbed (\pm 5%) Crustal thickness – varied (\pm 1.5 σ) Upper mantle velocity – perturbed (\pm 3%) Lithospheric thickness – varied (\pm 1.5 σ) Transitioned into ak135 model v_p/v_s fixed in seds/crust from CRUST1.0 set for lid, astheno fixed from ak135 in mantle



The dispersion values predicted by each model are then calculated using MINOS and compared to the observed dispersion from our high-resolution global surface wave model.

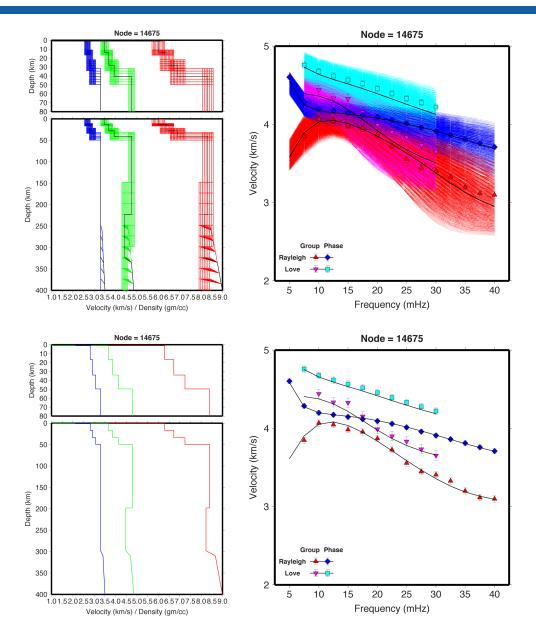
We select the model that fits the observed dispersion the best

Example – node 14675

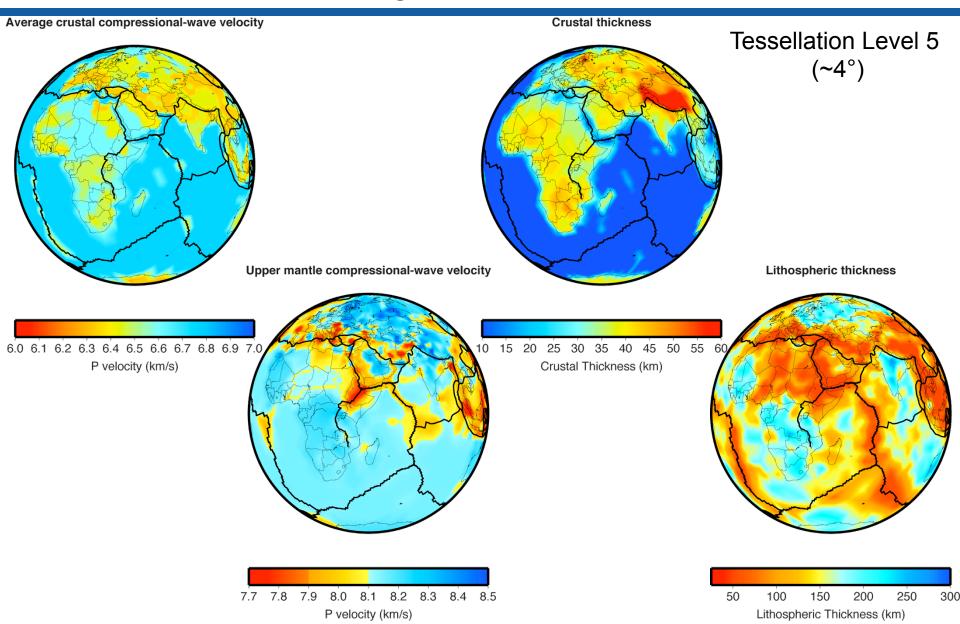


The range of models covers the parameter space in both depth space and dispersion space

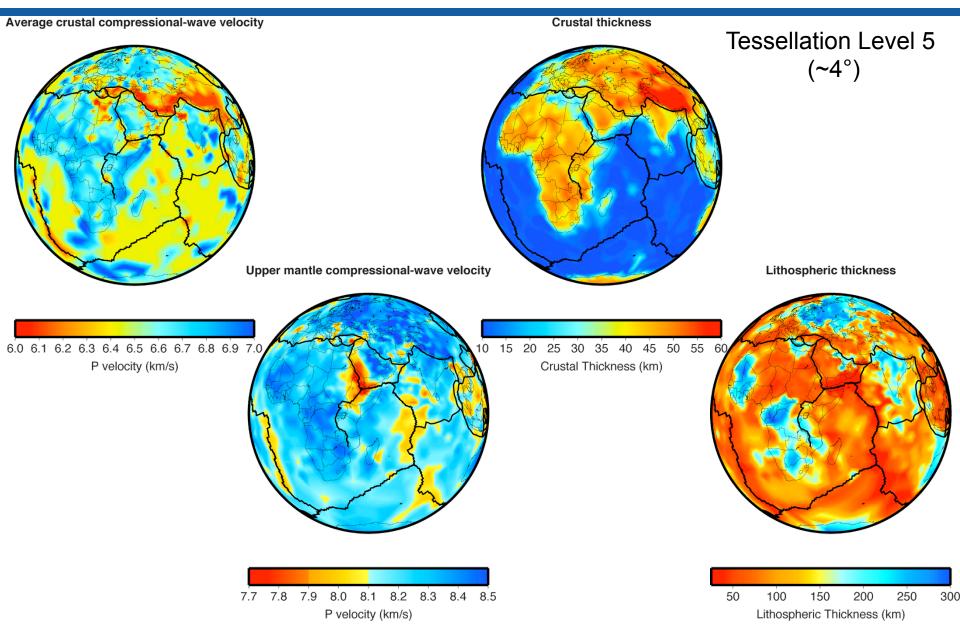
For this node, we are able to fit the surface waves with a faster crust, a slightly faster mantle, a thicker crust, and a thinner lithospheric thickness



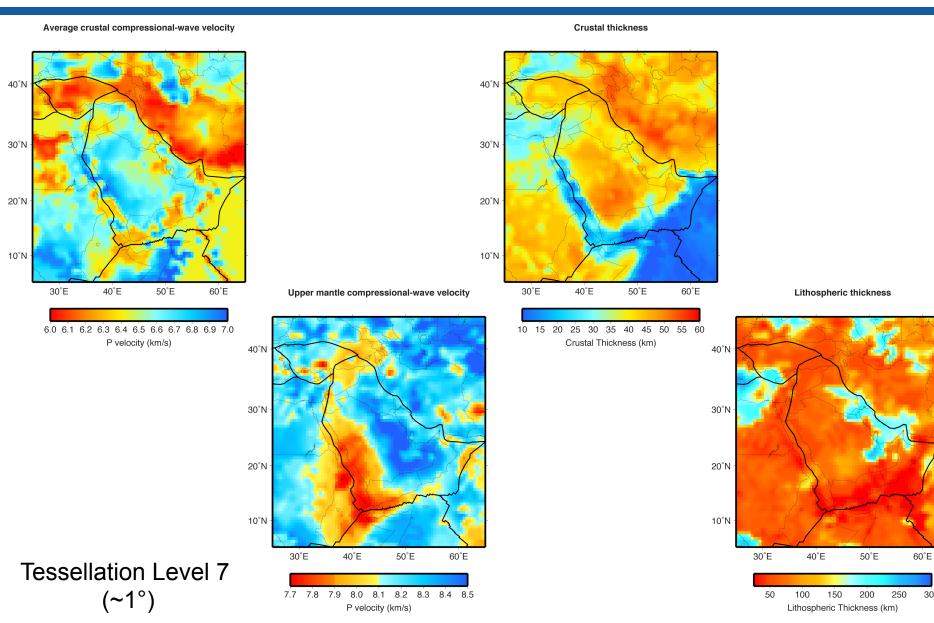
We assemble the results to construct the full model – Global, low-resolution, starting model



We assemble the results to construct the full model – Global, low-resolution, inverted model

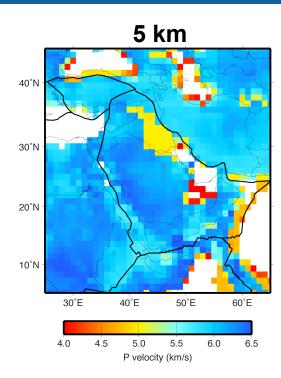


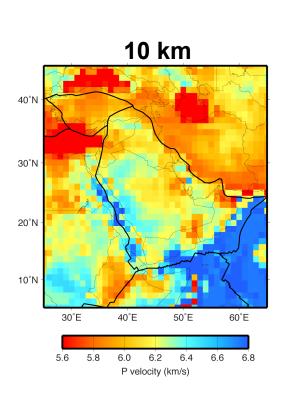
High-resolution regional results for the Middle East

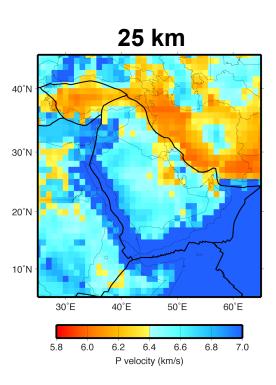


300

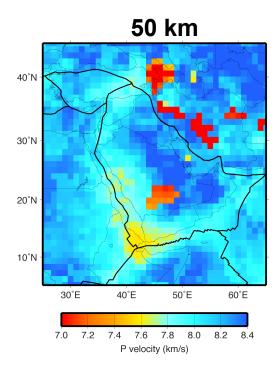
Depth slices through the model

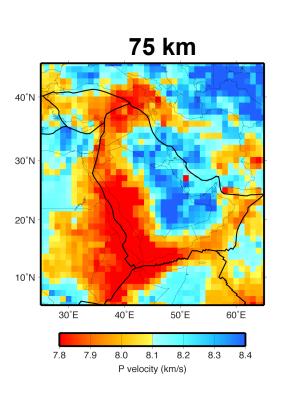


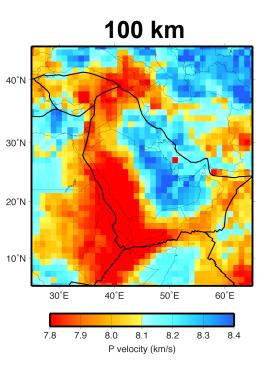




Depth slices through the model

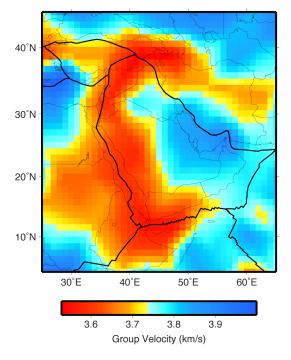




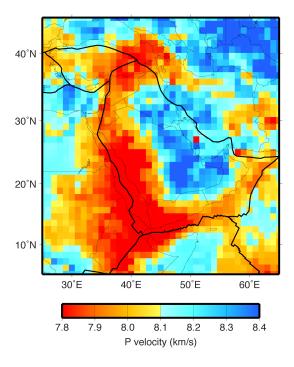


Close relationship between surface wave model and LITHO1.0

100 s Rayleigh wave group vel

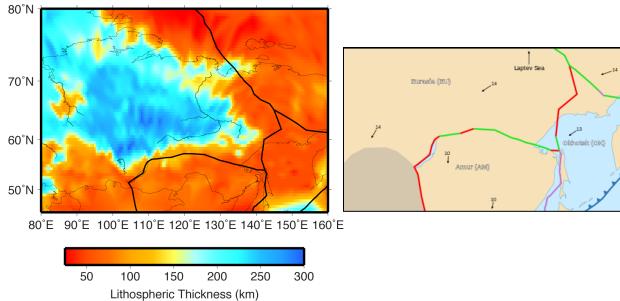


75 km depth slice



High-resolution regional results for Siberia

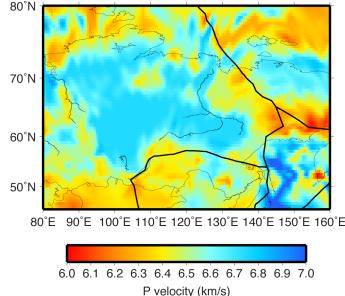
Lithospheric thickness



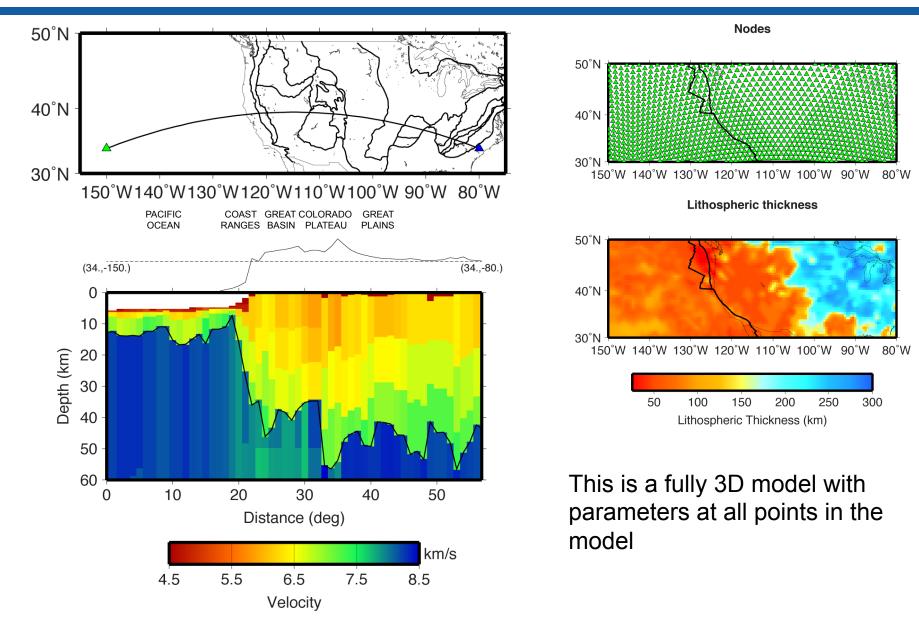
Constraints on lithospheric thickness can be used to provide information on the Siberian Shield and study the nature of this diffuse plate boundary between the Eurasia and North American Plates and the interaction with the Amur and Okhotsk Plates

Tessellation Level 7 (~1°)

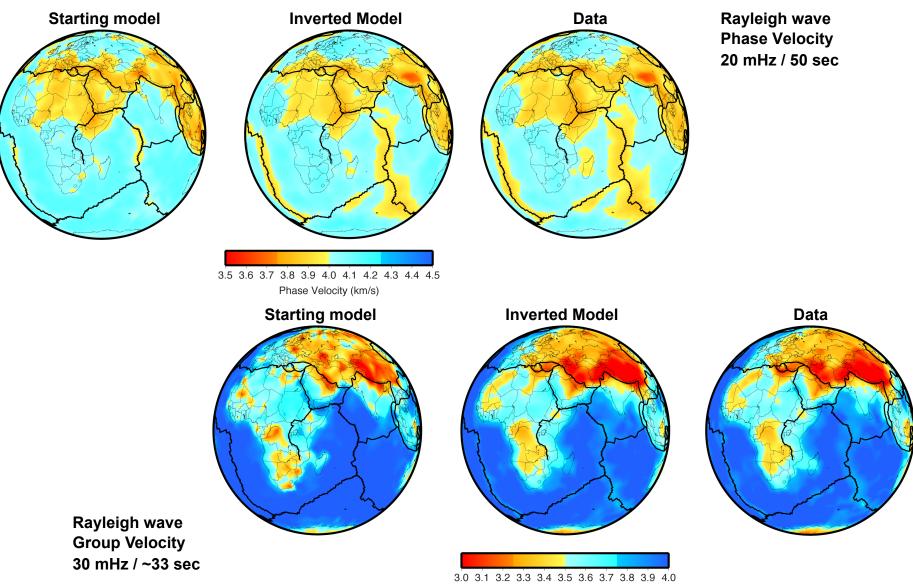
Average crustal compressional-wave velocity



High-resolution regional results for North America



The resulting model recovers the dispersion signal over a wide frequency range



Group Velocity (km/s)

The LITHO1.0 model appears to both fit the surface wave data and be consistent with other geophysical information and tectonic structure

We would like to validate this more rigorously with travel time, waveform, and other data

Finish the model

- We may still need to tweak the inversion to make crustal thickness changes smaller or convince ourselves that they are real
- We will incorporate anisotropy by allowing for a transversely isotropic mantle (in lid and asthenosphere)
- Make the runs at Tessellation Level 7 (~ 1°) globally

Make the model and interfaces available

- Create model in other formats (e.g. regular lat/lon grid, spherical harmonics?) from native tessellation format to be usable to the widest array of users
- Supply interfaces to provide depth profiles at arbitrary lat/lon locations and parameters (e.g. V_P, V_S, etc.) at arbitrary lat/lon/depth points.